Equation of state of supercooled water from the sedimentation profile

Masako Yamada, H. Eugene Stanley, and Francesco Sciortino

Center for Polymer Studies and Department of Physics, Boston University, Boston, Massachusetts 02215

Dipartimento di Fisica, INFM UdR and INFM Center for Statistical Mechanics and Complexity, Università di Roma “La Sapienza,” Piazzale Aldo Moro 2, I-00185 Roma, Italy

Received 23 August 2002; published 30 January 2003

To study the coexistence of two liquid states of water within one simulation box, we implement an equilibrium sedimentation method—which involves applying a gravitational field to the system and measuring or calculating the resulting density profile in equilibrium. We simulate a system of particles interacting via the Stillinger-2 (ST2) potential, a model for water. We detect the coexistence of two liquid phases at low temperature.

I. INTRODUCTION

The physics of the liquid state has been the subject of intense research activity. Novel approaches and novel apparatus have made possible the study of liquids under extreme thermodynamic conditions and in wide windows of space and time. Interesting and unexpected new phenomena have emerged as a result of a combined effort involving experiments, theory and simulations [1,2]. One of these is the possibility of a liquid-liquid (LL) transition in one-component systems—in addition to the usual liquid-gas transition. Several liquids [3–5] and several models [6–13] have been studied in detail, and it appears that the class of materials where a LL transition can be observed is larger than the class of tetrahedral liquids that were originally considered [14] possible candidates for a LL transition.

Water is one of the liquids that might possess a LL transition. Indeed, the first conjecture of a LL transition was based on a numerical study of the (ST2) potential [15], a model designed to mimic the behavior of liquid water. In the case of ST2, the two coexisting phases differ in their local structure. The low-density phase is formed by an open tetrahedrally coordinated network of hydrogen bonds, while the high-density phase has a more distorted network of hydrogen bonds. Recent theoretical work has shown that the interplay between local energy, entropy, and volume which may generate a LL transition can in principle be realized by spherically symmetric potentials [8,11].

The evaluation of the $P(\rho,T)$ equation of state (EOS) is key to test for phase coexistence (gas-liquid, liquid crystal, and LL), where $P$, $\rho$, and $T$ denote the pressure, density, and temperature. The numerical calculation of $P(\rho,T)$ requires the study of the model for a variety of state points. The coexistence between two phases, in the appropriate temperature window, appears as a region of density values where $P$ is a constant. In small-size numerical simulations it is sometimes hard to observe phase transitions directly in one simulation box, in part because the free energy associated with creating an interface often stabilizes metastable phases [16]. In these cases, $P(\rho,T)$ does not show any flat region [17].

A different approach for studying in one single numerical simulation an entire isotherm has been proposed in Ref. [18] (and later exploited in the experimental study of colloidal systems and in the study of crystallization profiles Refs. [19–22]). This approach simulates a semi-infinite tube in the presence of a very strong gravitational field and measures the density profile in equilibrium. A simple inversion of the density profile allows the model EOS to be constructed. This idea has also been applied in the experimental study of the EOS of colloidal particles, by inversion of the measured sedimentation equilibrium profile.

Here, we apply the sedimentation profile method to ST2 water. We find that an interface separating two liquid states appears at low temperature, corresponding to the coexistence of two metastable liquid states of water within one simulation box, providing evidence for the presence of a LL transition.

II. THEORY AND SIMULATION DETAILS

We study $N=7680$ rigid molecules of mass $M$ interacting through the ST2 water potential, a rigid, nonpolarizable, five-site potential [23] that is able to reproduce qualitatively the thermodynamic anomalies of liquid water [24]. “ST2 water” is characterized, on cooling, by isobaric density maxima, increasing compressibility, increasing constant $P$ specific heat, and evidence for a LL transition in the deeply supercooled regime (which is difficult to probe experimentally due to spontaneous crystallization [15]).

To implement the sedimentation profile method, we use a column-shaped simulation box, semi-infinite along the $z$ axis and with periodic boundary conditions along the $x$ and $y$ axes. The top of the box is left open, while the bottom is assumed to be a repulsive soft-sphere surface, generating a short-range force proportional to $z^{-13}$ acting on the molecule center of mass (Fig. 1). The box width in the $x$ axis and $y$ axis directions is $3\,\text{nm}$, corresponding to a bottom surface area of $S_{xy}=9\,\text{nm}^2$. A strong gravitational field $g_s=2\times10^{12}\,\text{g}$ is applied downward, in the $z$ axis direction, where $g=9.8\,\text{kg/m}^2$ is the Earth’s gravitational field. The value of the field controls the range of $P$ values accessed in the simulations. The pressure at the bottom of the column is $P=g_NVMS_{xy}=500\,\text{MPa}$ [25].

The chosen gravitational field $g_s$ is able to generate a pressure difference of about 500 MPa over a distance of the order of 100 molecular diameters. Compared to a standard molecular dynamics (MD) simulation, each molecule is subject to an additional force in the $z$ direction arising from the
gravitational field which, in the present case, is $10^{-4}$ times less than the average intermolecular force. In other words, $M g_z/(\langle F_i^z \rangle) \sim 10^{-4}$, where $\langle F_i^z \rangle$ is the absolute value of the $z$ component of the force on the center of mass of the $i$ molecule, averaged over all molecules. Thus, for each molecule, the gravitational force is negligible compared to the intermolecular interaction. The gravitational field acts only as a device to produce a density gradient in the box. It generates, in each slab of the simulation, conditions identical to the ones that would have been generated in a conventional simulation with the same average density.

We perform the simulations using a multiprocessor code on SGI Origin 2000, IBM SP, and IBM Regatta supercomputers. We choose a 1 fs time step and study different temperatures from $T = 300$ K down to $T = 230$ K. For $T = 230$ K, we simulate four different systems to better estimate error and reproducibility of the results. A long ($\sim 10$ ns) equilibration time precedes the actual calculation of the equilibrium density profile. To analyze equilibration, we monitored the running average of the $z$ position of the center of mass. Production runs lasted at least 200 ps for the higher temperatures up to several nanoseconds per box, when $T = 230$ K. The total simulation time in each of our calculations does not exceed 20 ns. This time is sufficient to guarantee proper equilibration of the metastable (as compared to the crystal) liquid and the evaluation of “equilibrium” averages in the supercooled liquid phase. The formation of a crystalline phase due to homogeneous nucleation requires much longer simulation times and it is never observed either in our calculations or in conventional simulations. In this respect, the time scale of simulations offers the possibility to look into a temperature region where experimental observation of the liquid phase is hindered by homogeneous nucleation.

We define a mass density field $\rho(z)$ by averaging the density over bins with height $\Delta z = 1$ nm to calculate the pressure field $P(z)$ from the relation $P(z) = g \rho(z) \Delta z$. A parametric plot of $P(z)$ vs $\rho(z)$ provides the EOS.

III. RESULTS AND DISCUSSION

Figure 2 shows the equilibrium density profile along the $z$ axis for three different temperatures. Each symbol represents one bin, where the height of each bin $\Delta z = 1$ nm. For $T = 300$ K, we see only one break in the density profile: the topmost points in the plots correspond to the gas-liquid interface. Above these points is a much less dense gas not shown on this density scale. For $T = 250$ K, there appears an inflection. For $T = 230$ K, we see a clear break in the density, associated with the interface between two different liquid states.

Figure 3 shows the corresponding $P(\rho)$ relations at the same three temperatures. We also show $P(\rho)$ as evaluated previously using standard MD for cubic boxes with periodic boundary conditions for systems with $N = 6^3 = 216$ (Ref. [26]) and $N = 12^3 = 1728$ (Ref. [27]). In Refs. [26,27] $P$ was calculated with the standard virial relation [28].

Consistent with the data shown in Fig. 2, we see a region of coexistence between two different phases at $T = 230$ K, in agreement with the estimated location of the LL transition in Ref. [29]. Note that when using the sedimentation equilibration method, we see no unphysical loops in the equation of state, unlike cubic-box simulations where the boundary conditions may artificially stabilize metastable states.

To confirm that the two coexisting phases are both liquids, we calculate the mean square displacement for different height values, both below and above the interface. We follow the evolution of each molecule for an average mean squared
displacement smaller than $D_z$, so that each height value can be unambiguously assigned to an average density value. Figure 4 shows that both phases are sufficiently diffusive. The low-density phase has smaller diffusivity, in agreement with previous simulations of the density dependence of the dynamics.

IV. CONCLUSIONS

We studied the EOS of “ST2 water” with the sedimentation profile method, and presented evidence that the EOS at $T=230$ K shows a clear phase coexistence, between two phases which are both liquids. The observed LL coexistence in the ST2 potential phase diagram is consistent with estimate of the “critical point” being located around $T_c=235$ K, $P_c=250$ MPa, and $\rho_c=1.05$ g/cm$^3$. Note that for most simple water models, the temperature and pressure scales are shifted relative to real values, thus, they must be shifted to place them within an experimental context.

Indirect experimental measurements are consistent with the possibility of a critical point with coordinates around $T_c=230$ K, $P_c=100$ MPa, and $\rho_c=1.05$ g/cm$^3$.

Note added in proof. Very recently, I. Brovchenko, A. Geiger, and A. Oleinikova (preprint) found three distinct liquid-liquid phase transitions in ST2 water when the long-range interactions are treated by a simple spherical cutoff without reaction field, producing exactly the correct density at the liquid-vapor coexistence line.

ACKNOWLEDGMENTS

This work was supported by NSF Grant No. CHE-0096892. M.Y. was also supported by NSF Grant No. GER-9452651. F.S. acknowledges support from MIUR FIRB and COFIN and INFM-Initiative Parallel Computing for granting numerical resources.
We use the standard smoothed cutoff of 2.5σ + 0.005 nm, where σ (0.31 nm) is the characteristic radius of the ST2 water molecule, given by the Lennard-Jones component of the potential. We include long-range corrections using the reaction field method.

[25] We performed a numerical check of the simulation results to confirm that the strong external field does not affect the intermolecular interactions among the water molecules. Using the same molecular configurations, we compared the equation of state calculated using the “standard” virial expansion method (where intermolecular forces between all molecule pairs are explicitly taken into account) and the equation of state calculated using the present method. The results are identical, giving confidence that the results obtained in such a large gravitation field are equivalent to results obtained at huge pressures. To intuitively understand the degree to which increasing g corresponds to increasing pressure, consider that pressure at given depth of water is generated by the force of gravity on molecules above the sample, so 100 MPa pressure can be achieved by submerging a sample to a depth of 10,000 m on earth, to 4000 m on Jupiter (where g = 2.5 g), or to only 10⁻⁸ m = 10 nm if g = 10⁻². 


[28] Regions of negative P cannot be studied with the sedimentation method, at least not in its present form.


[31] In comparison, the critical point estimated using the extended simple point change potential (SPC/E) is estimated to be $T = 130$ K and $P = 290$ MPa with $ρ = 1.10$ g/cm³ [33]; using the transferable intermolecular potential with 4 points (TIP4P) $T < 200$ K and $P > 70$ MPa [29]; and using the TIP5P potential $T_c ≈ 217$ K, $ρ_c ≈ 1.13$ g/cm³, and $P ≈ 340$ MPa [34].

